# Tsinghua Center for Astrophysics (Beijing)



# Signatures of Cosmic Reionization on the 21cm 3-Point Correlation

Kai Hoffmann

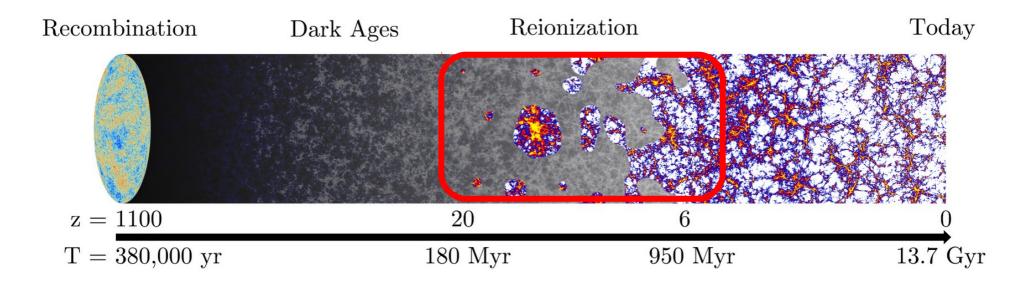
collaborators:

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### **Motivation**

#### 21cm clustering statistics can constrain

- Reionization models (Shimabukuro et al. 2017, 1608.00372)
- possibly cosmological models at high z (see conclusions)



Credit: Kaurov

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#### Why going beyond 2-point statistics?

21cm 2pc probes scale dependence of 21cm clustering, while 3pc is sensitive to additional information on shape of

ionized regions



tighter model constraints

matter fluctuations

(Suman Majumdar et al. 2017, arXiv:1708.08458)

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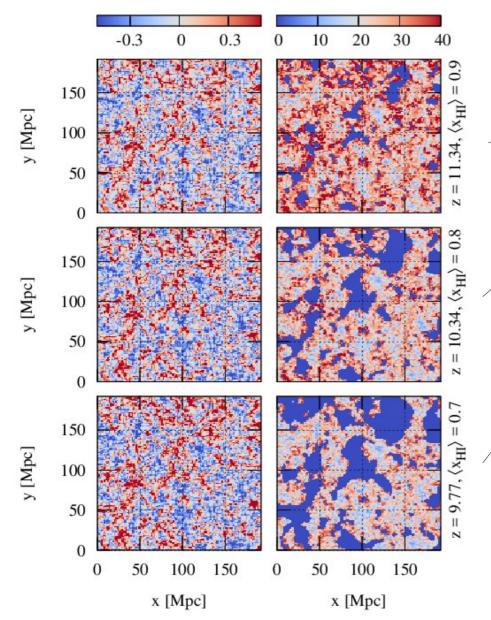
tighter model constraints

matter fluctuations

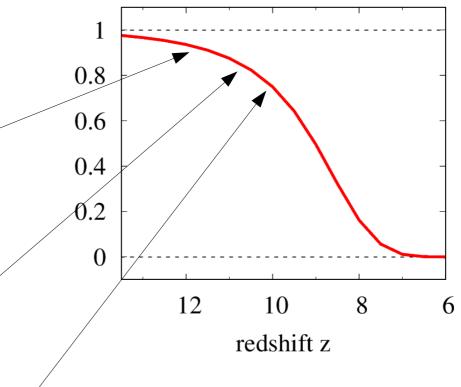
no theory model for the 21cm 3pc

### 21cmFast simulations





#### Global neutral fraction



• (768 Mpc)<sup>3</sup> box

200 realizations

(21cmFAST: Mesinger et al. 2010)

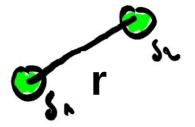
# 2- and 3-point correlations

$$\delta_{m} = \frac{\rho_{m} - \langle \rho_{m} \rangle}{\langle \rho_{m} \rangle}$$

fluctuations: 
$$\delta_m = \frac{\rho_m - \langle \rho_m \rangle}{\langle \rho_m \rangle}$$
  $\delta_{\delta T} = \frac{\delta T - \langle \delta T \rangle}{\langle \delta T \rangle}$ 

#### 2-point correlation (2pc):

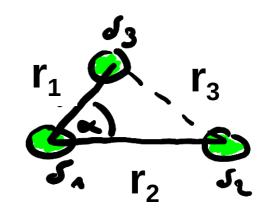
$$\xi^{(12)} \stackrel{\text{def}}{=} \langle \delta_1 \delta_2 \rangle (r)$$



• spherically symmetric → not sensitive to shape of fluctuations

#### 3-point correlation (3pc):

$$\zeta^{(123)} \stackrel{\text{\tiny def}}{=} \langle \delta_1 \delta_2 \delta_3 \rangle (r_1, r_2, r_3)$$

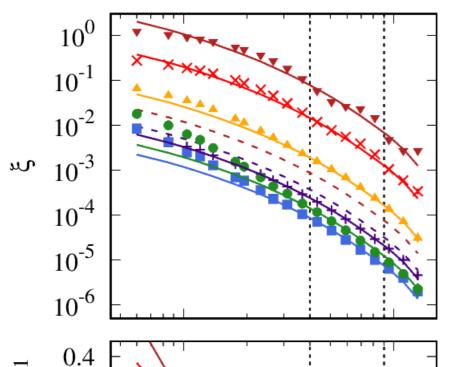


provides additional shape info

# **2pc measurements**

$$\langle x_{HI} \rangle = 0.99 + \langle x_{HI} \rangle = 0.80 + \langle x_{HI} \rangle = 0.95 + \langle x_{HI} \rangle = 0.60 + \langle x_{HI} \rangle = 0.90 + \langle x_{HI} \rangle = 0.30 + \langle x_$$

mean measurements over 200 realizations



10

r [Mpc]

100

0.2

-0.2

-0.4

0

symbols: 21cm 2pc ( $\xi_{8T}$ )

dashed lines: matter 2pc ( $\xi_m$ )

**solid lines:** fits to bias model

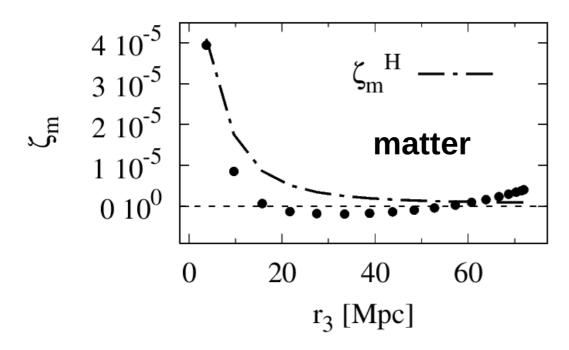
#### **Leading-order bias model**

$$\xi_{\delta T} = b_1^2 \xi_m$$

fitting range: 40 < r < 90 Mpc

# **3pc measurements**

$$(r_1,r_2) = (36,36)$$
 [Mpc]

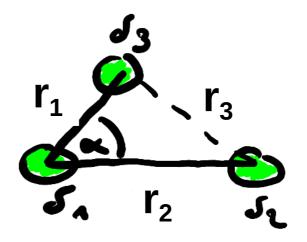


dots: 3pc

$$\zeta \stackrel{\text{def}}{=} \langle \delta_1 \delta_2 \delta_3 \rangle$$

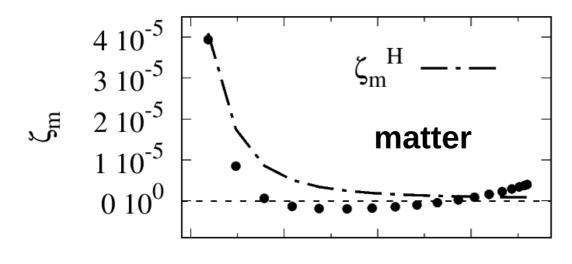
line: hierarchical 3pc

$$\zeta^{H} \stackrel{\text{def}}{=} (\xi_{m}^{(12)} \xi_{m}^{(13)} + 2 perm.)$$



# **3pc measurements**

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 [Mpc]

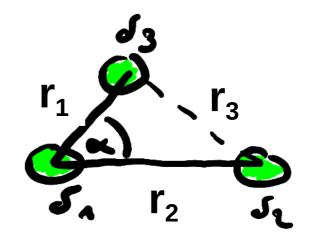


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# quadratic bias model

**Assumption:** 

21cm brightness temperature is deterministic function of underlying matter density

$$\delta_{\delta T} = F(\delta_m) \approx \sum_{n=0}^{N} b_n \frac{\delta_m^n}{n!}$$

(Taylor expansion of F around  $\delta_m \approx 0$  )

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Leading-order approximations for biased correlation functions

2pc 
$$\xi_{\delta T} \approx b_1^2 \xi_m$$

$$3pc \qquad \zeta_{\delta T} \approx b_1^3 \zeta_m + b_1^2 b_2 \zeta_m^H$$

(Fry & Gaztanaga 93)

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Leading-order approximations for biase

$$2pc \quad \xi_{\delta T} \approx b_1^2 \xi_m$$

(Taylor expansion of F around  $Q_{\delta T} \approx \frac{1}{b_1} (Q_m + \frac{b_2}{b_1})$  approximations for biased with  $Q \stackrel{\text{def}}{=} \zeta/\zeta^H$ 

with 
$$Q \stackrel{ ext{def}}{=} \zeta/\zeta^H$$

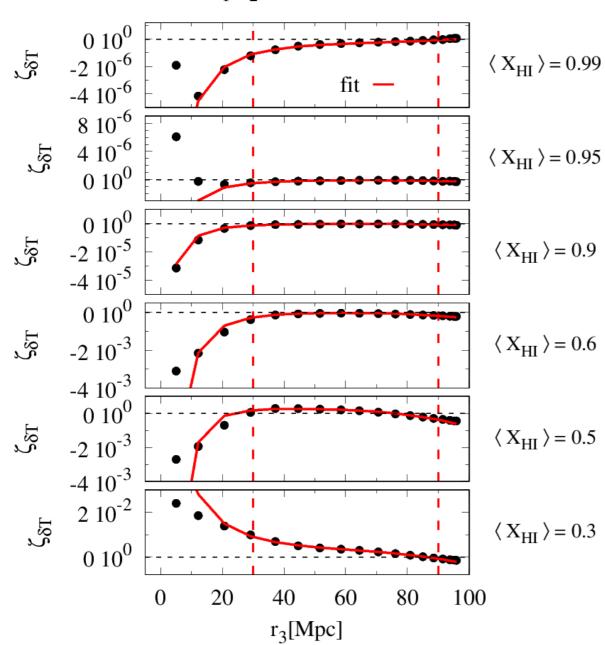


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# 21cm 3pc fits

$$(r_1, r_2) = (48, 48) \text{ Mpc}$$



dots: 21 3pc

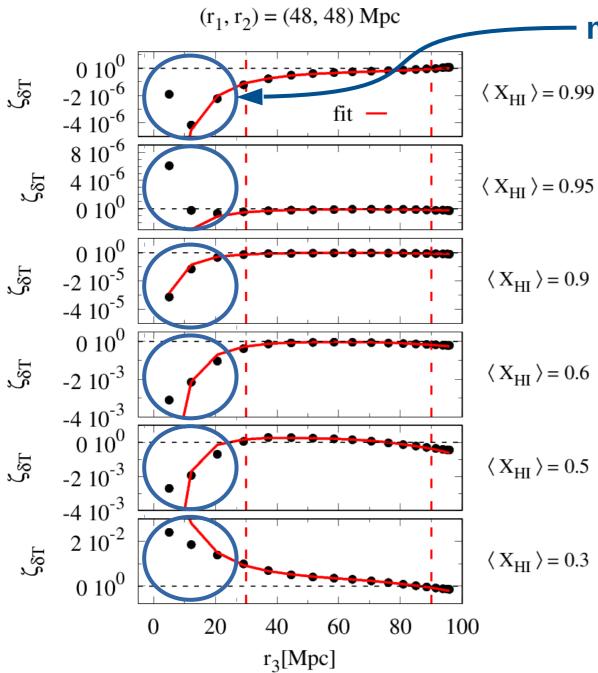
$$\zeta \stackrel{\text{def}}{=} \langle \delta_1 \delta_2 \delta_3 \rangle$$

lines: 3pc bias model fit

$$\zeta_{\delta T} \approx b_1^3 \zeta_m + b_1^2 b_2 \zeta_m^H$$

(fitting range: 30<r<90 Mpc)

# 21cm 3pc fits



#### model fails at r3 < 20 Mpc

dots: 21 3pc

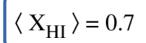
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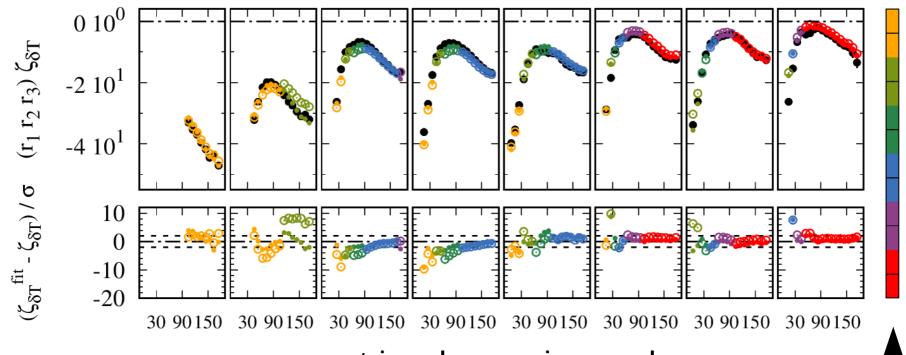
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# 21cm 3pc fits



triangle configurations, defined by (r1, r2)

(36,24) (48,24) (48,48) (54,42) (60,36) (60,60) (72,48) (72,72)



triangle opening angle

• black dots: 3pc measurements

• colored dots: fits to bias model prediction in triangle scale bins

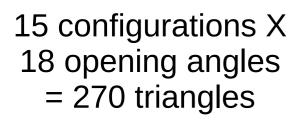
37.6 44.4 50.7 57.6 64.5 77.4

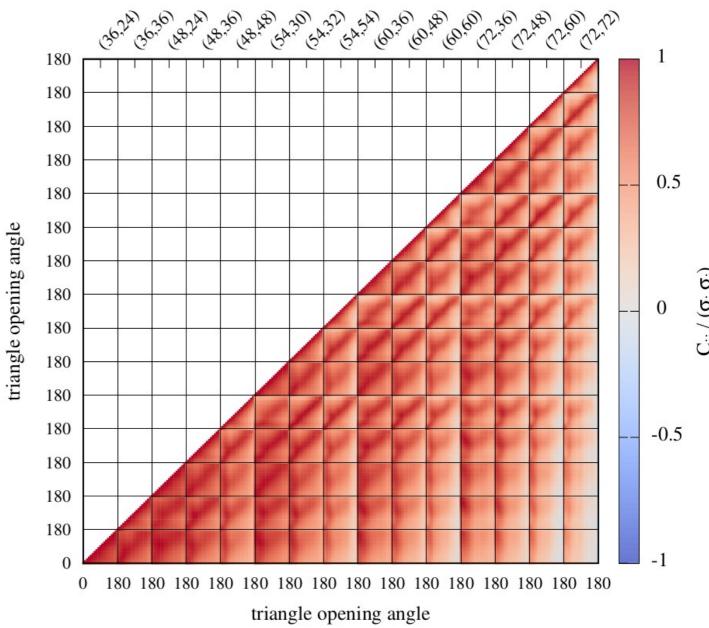


$$(r_1 r_2 r_3)^{1/3}$$

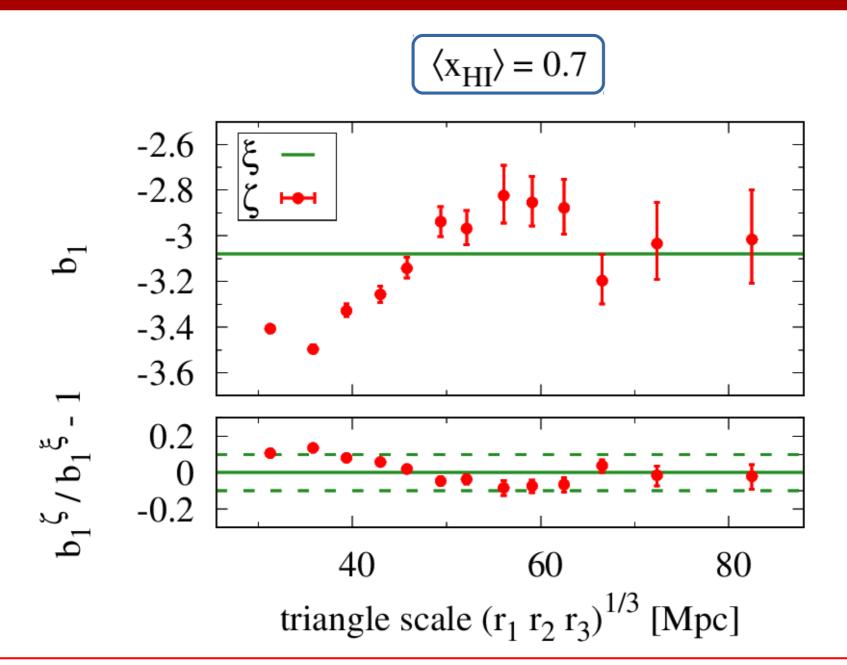
# 21cm 3pc covariance

triangle configuration



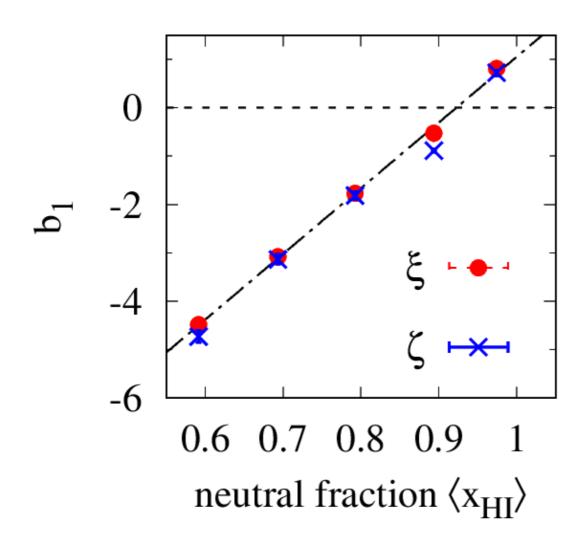


### linear bias measurements



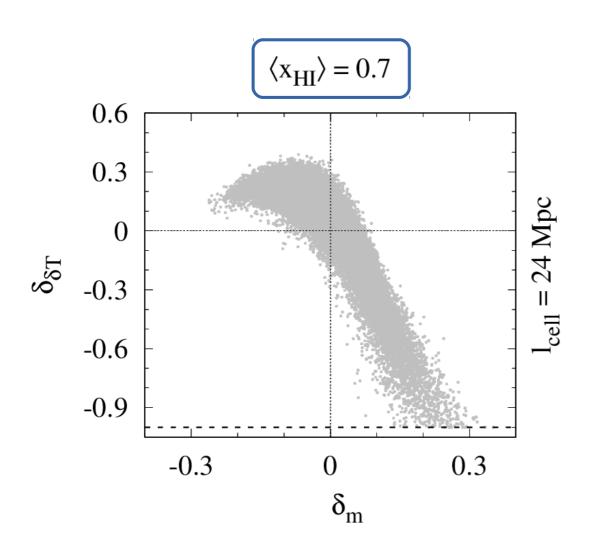
~10% agreement between linear bias from 2pc and 3pc

# Linear bias comparison with 2pc

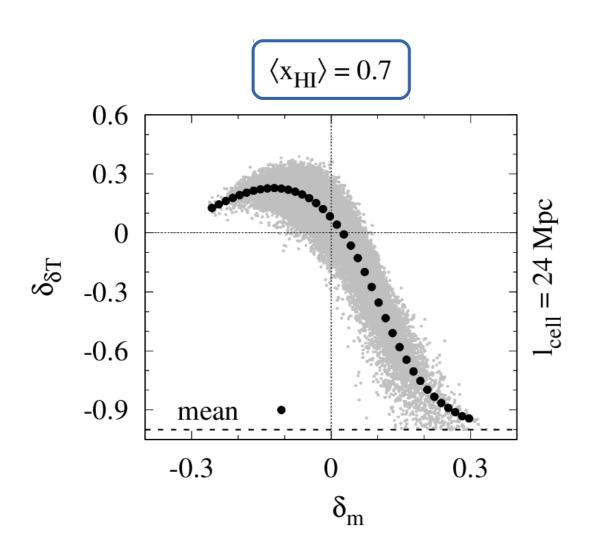


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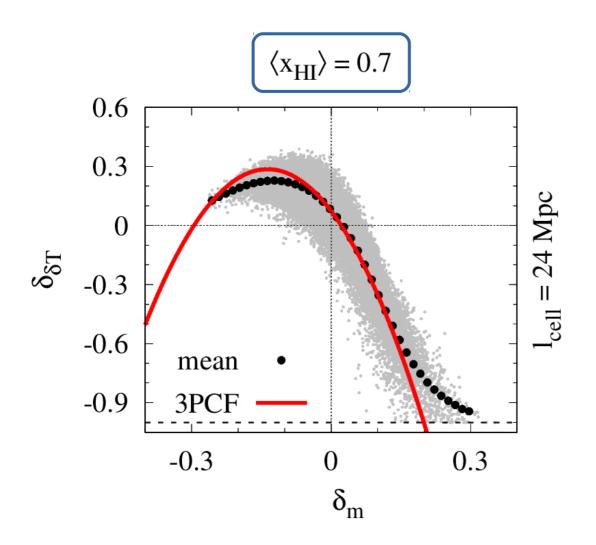
$$\delta_{\delta T} = F(\delta_m)$$
?



 grey dots: fluctuations in 24 Mpc cubical grid cells in one realization



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- black dots: mean  $\delta_{\delta T}$  over 200 realizations in  $\delta_m$  bins



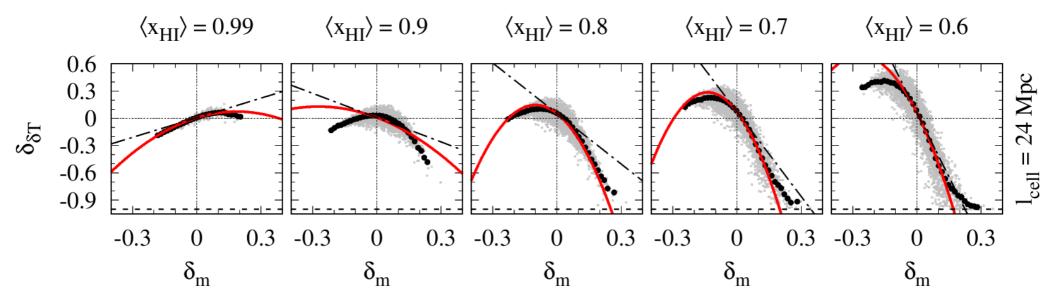
- grey dots: fluctuations in 24 Mpc cubical grid cells in one realization
- black dots: mean  $\delta_{\delta T}$  over 200 realizations in  $\delta_m$  bins
- red line: quadratic bias model with b1 and b2 from 3pc measurements

$$\delta_{\delta T} = b_0 + b_1 \delta_m + b_2 \delta_m^2$$

$$(\langle \delta_{\delta T} \rangle = 0 \rightarrow b_0 = -b_2 \langle \delta_m^2 \rangle)$$

quadratic bias model with b1 and b2 from 21cm 3pc:

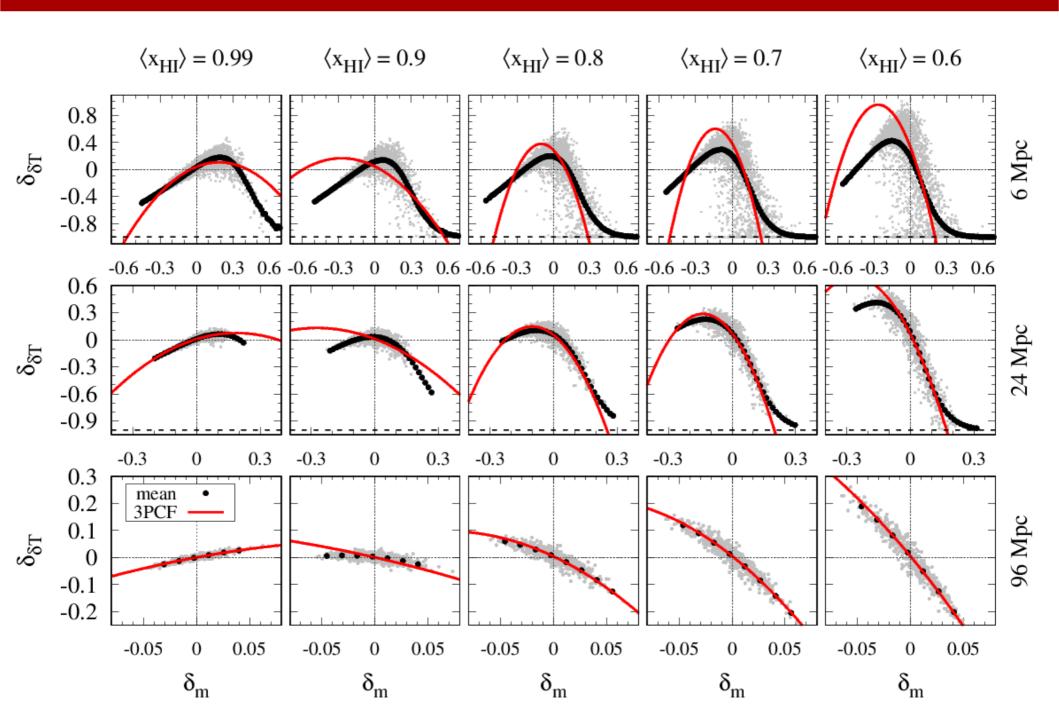
$$\delta_{\delta T} = b_1 \delta_m + b_2 (\delta_m^2 - \sigma_m^2)$$

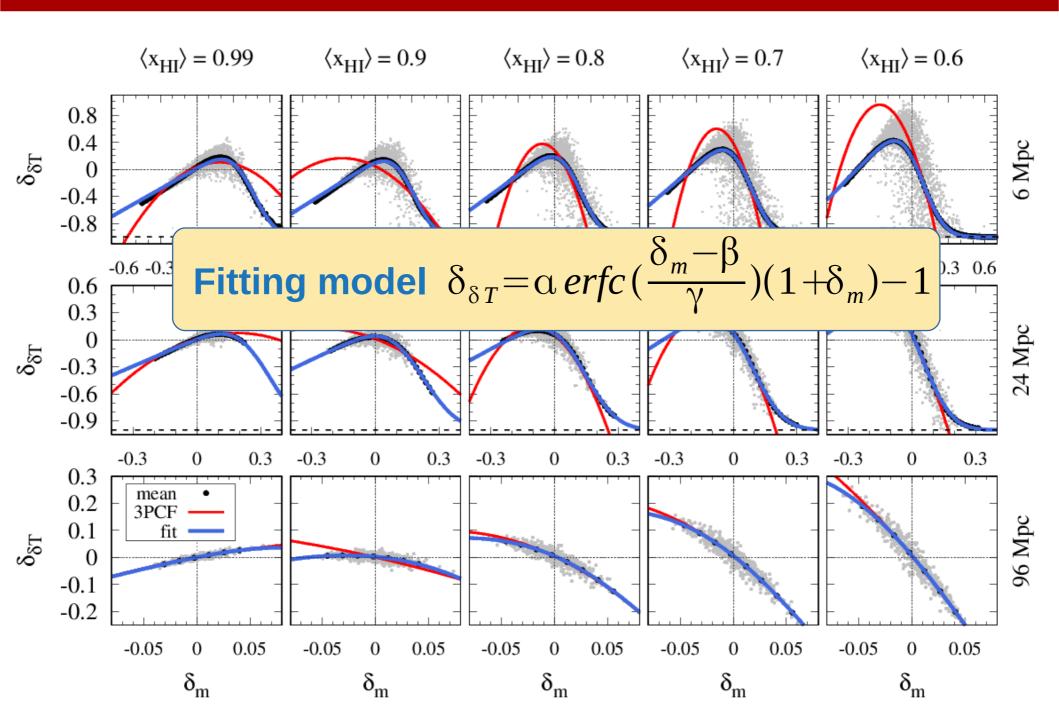


**b1** = slope of  $\delta_{\delta T}$  at  $\delta_m = 0$ 

- positive at early times
- negative at late times

$$\delta T \propto \rho_{HI} \propto x_{HI} (1 + \delta_m)$$





### conclusions

- Quadratic bias model explains shape of 21cm 2pc & 3pc at large scales and early times of reionization (neutral fraction > 60%, r> 20 Mpc)
- b1 from 2pc and 3pc consistent at 10% level
- b1 and b2 measurements might allow for extracting physical information on EoR from 21cm observations
- Combining 21cm 2pc & 3pc can break growth-bias degeneracy

$$rac{\xi_{\delta T}(z)}{\xi_{\delta T}(z_0)} \propto rac{D(z_0)^2}{D(z)^2} rac{b_1(z_0)^2}{b_1(z)^2}$$

→ cosmological constraints from growth measurements at high z

arXiv:1802.02578

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